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The ultra sensitive GOCE Accelerometers and their future developments

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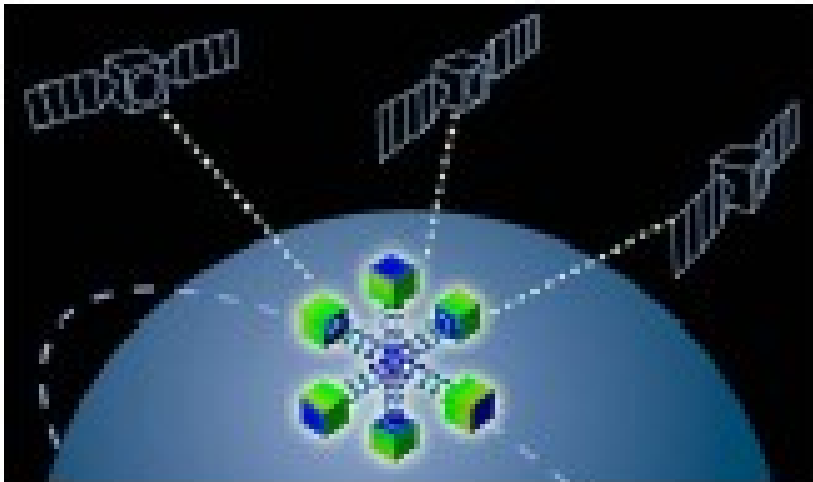
Towards a Roadmap for Future Satellite Gravity Missions

GRAZ – AUSTRIA – September 30- October 02 2009



r e t u r n o n i n n o v a t i o n

The realization of the accelerometers of the  *GOCE mission* was carried out under contract with  acting as *Prime Contractor of the Satellite*.



Acquisition Mode:

April 7th, 9h06 UTC

6 Proof masses in levitation

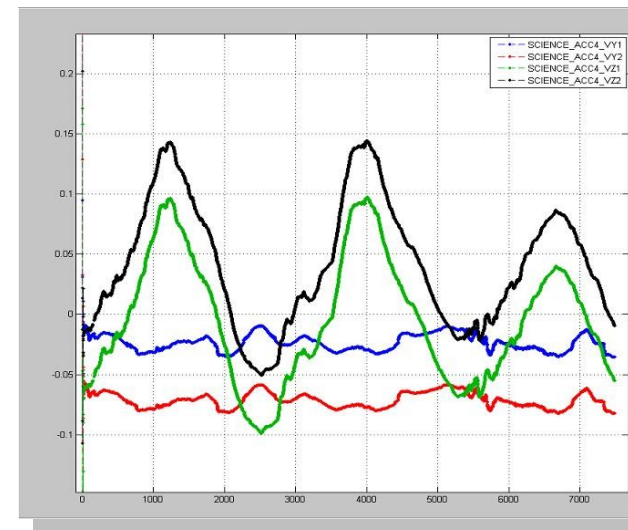
Science Mode:

April 8th, 8h02 UTC

Launch on March 17th



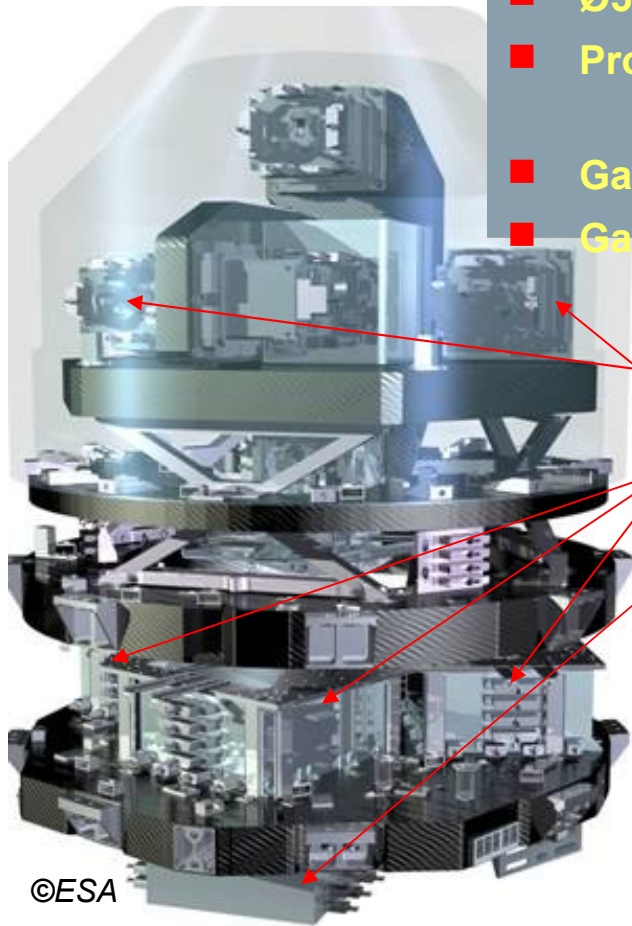
©ESA



Before to come to such a signal

GRADIOMETER / ACCELEROMETER CONFIGURATION

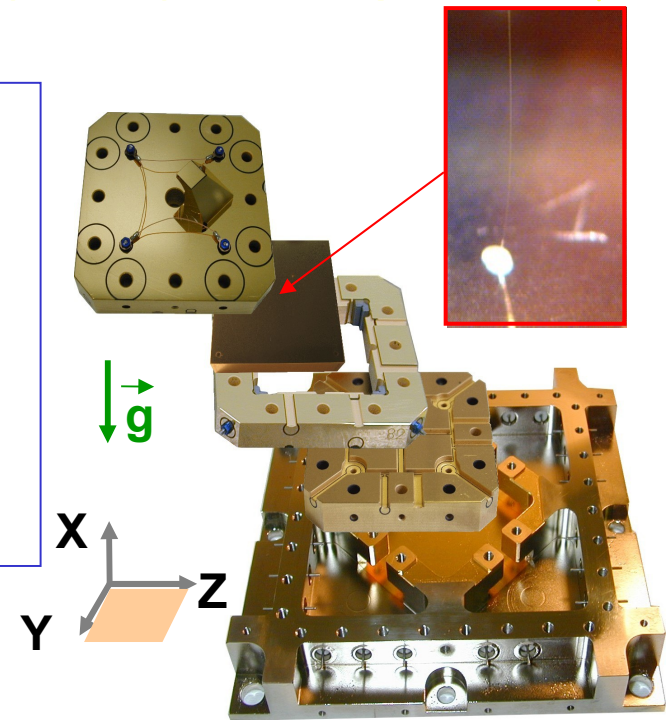
- Ø5 µm Goldwire for PM DC voltage bias
- Proof mass : 320g platinum-rhodium alloy (40 x 40 x 9.992 mm)
- Gap X : proof mass - the 4 x electrode pairs = 32 µm
- Gap YZ : proof mass - the 4 (Y and Z) electrode pairs = 299 µm



- 3 pairs of Accelerometer Sensor Head (ASH)
- 3 Front End Electronics Unit
- 1 Gradiometer Interface Unit
- 1 Carbon-Carbon stable structure
- 3 stages accurate thermal control

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GOCE Gradiometer



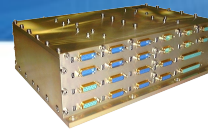
ASH mechanical core

CONTROL LOOP AND DATA PROCESSING

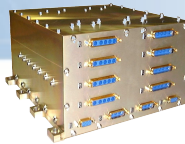


ASH

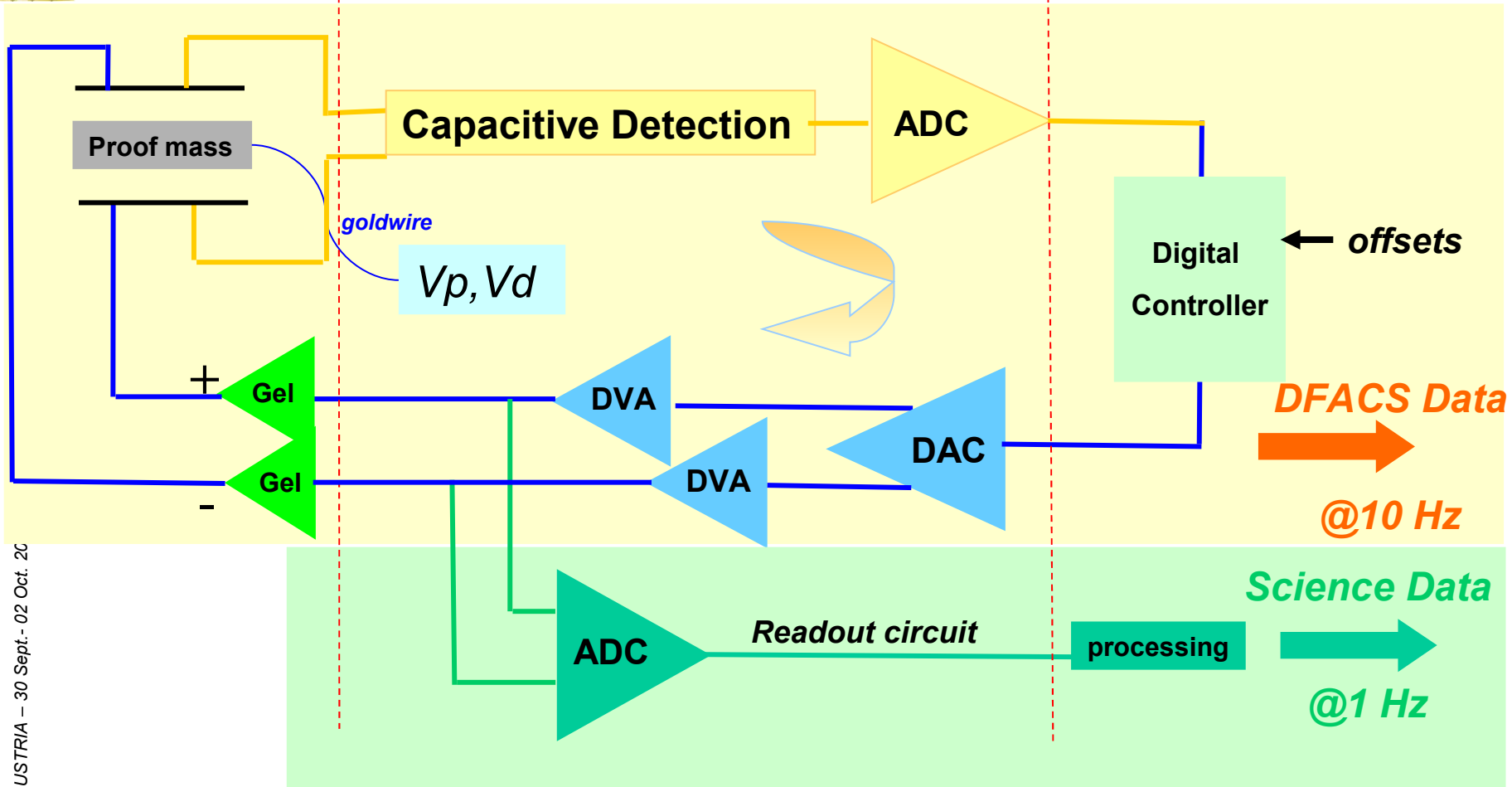
FEEU



GAIEU



Roadmap for Future Satellite Gravity Missions
GRAZ - AUSTRIA - 30 Sept. - 02 Oct. 20



$$\Gamma(V_{sc}) = bias + G(1+K_1) V_{sc} + K_2 G^2 V_{sc}^2 + noise$$

ACCELEROMETER PERFORMANCES (Science mode)

PM Polarisation Voltage

$$V_p = 7.5 \text{ V}$$

Detector gain $1.7 \text{ mV} / \text{nano-m}$

Scale factor

Science data 100 nano-g/V

DFACS data 1.7 micro-g/V

Resolution $< 2 \cdot 10^{-12} \text{ ms}^{-2} \text{ Hz}^{-1/2}$

Range $\pm 6.5 \cdot 10^{-6} \text{ ms}^{-2}$

Main contributors in MBW

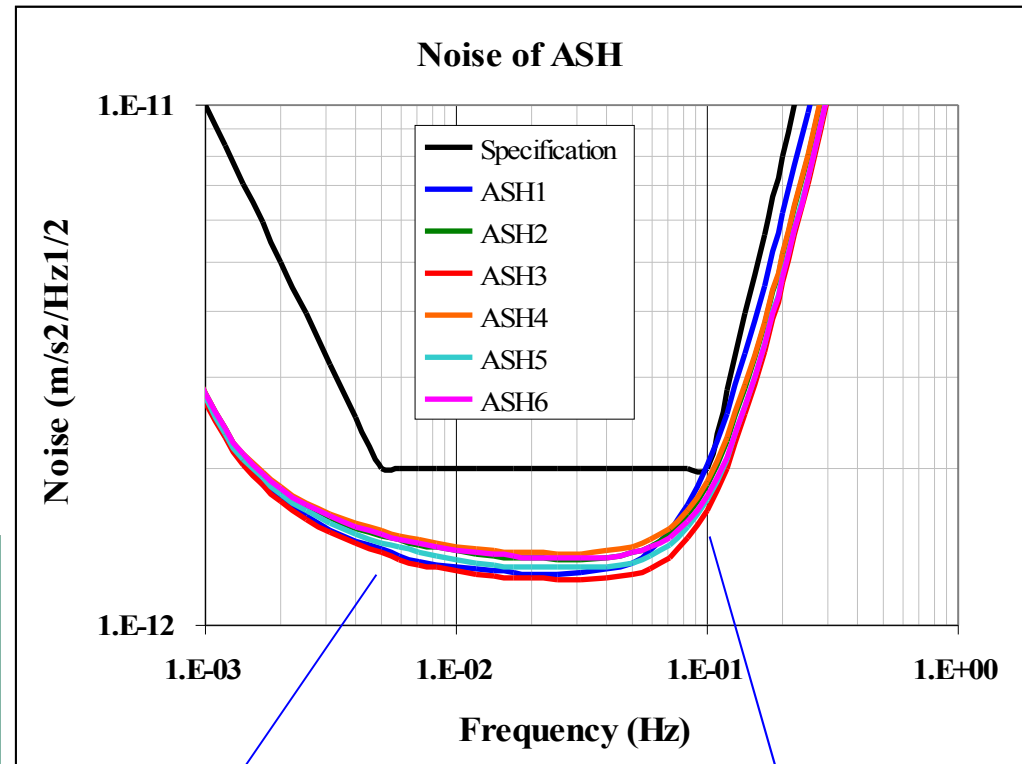
Upper Part : Detector-ADC1, ADC 2

Lower Part : ADC2, Goldwire damping

Bias thermal sensitivity

PM Detection Voltage

$$V_d = 7.6 \text{ V @ } 100 \text{ KHz}$$



$$1.38 - 1.52 \times 10^{-12} \text{ ms}^{-2} \text{ Hz}^{-1/2} \text{ @ } 5 \text{ mHz}$$

$$1.67 - 2.02 \times 10^{-12} \text{ ms}^{-2} \text{ Hz}^{-1/2} \text{ @ } 100 \text{ mHz}$$


TEST PLAN : from parts to Unit

At Parts level:

- *Dedicated Physics experiments : tribology, CPD, stiffness...*
- *Accurate geometrical control: down to $1\mu\text{m}$ for core parts*
- *Test bench for Electronics performance and thermal sensitivity : down to $0.1\mu\text{V}$ and 10^{-15} F*

At Unit level, on ground tests need:

- *to levitate the PM against gravity*
- *a low acceleration environment to not saturate electronics*

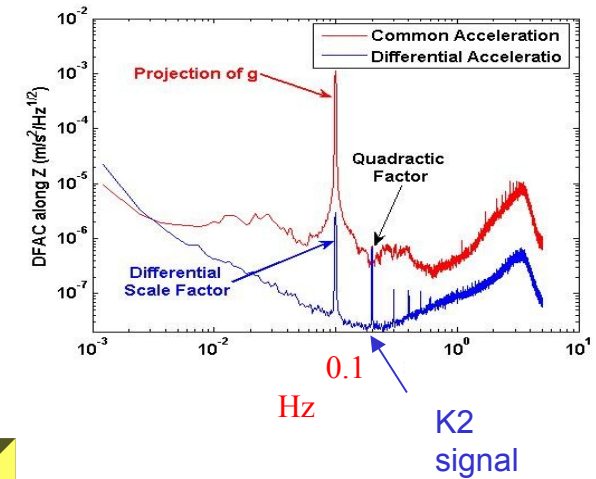
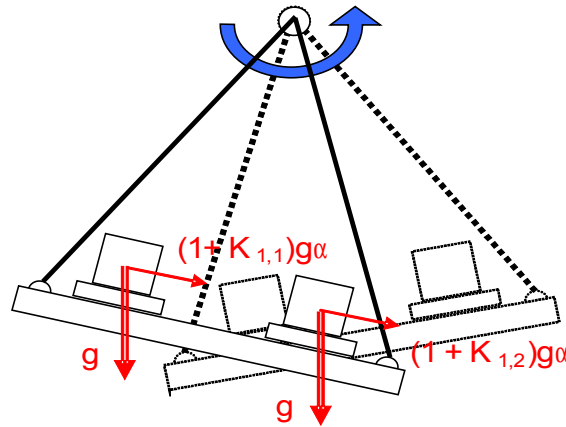
- 
- ☐ *ASH X axis dedicated to 1 g levitation of the PM*
 - ☐ *Pendulum bench controlled in horizontality better than $1\mu\text{rad}$*
 - ☐ *Free Fall test in low gravity conditions*

Application to:

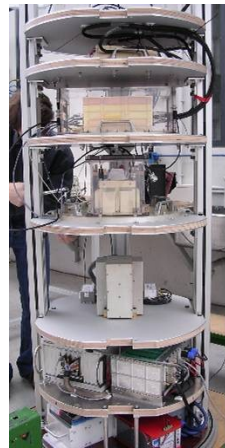
- *Functional test and transfer function, linearity, range, stiffness verification,...*
- *Scale factor and Quadratic factor verification,*
- *FDIR software verification,*

ON GROUND TESTS

Creation of a controlled in-plane acceleration by tilting the pendulum



Free Fall Test in ZARm drop tower



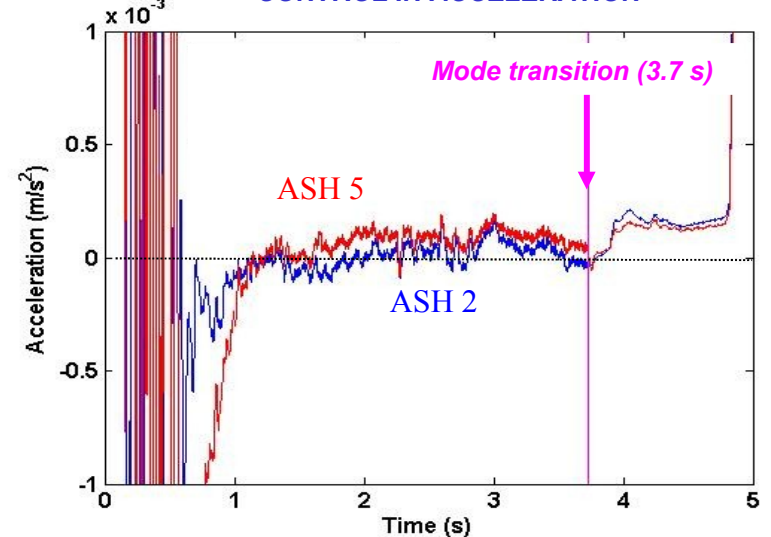
FEEU

ASH Pair

Super STAR Accelerometer

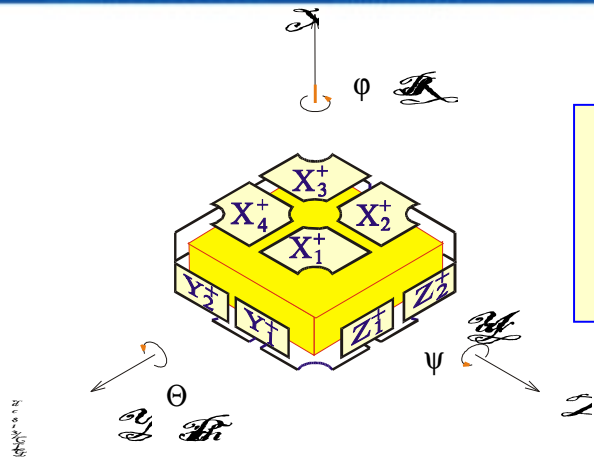
Capsule computer & batteries

**X AXIS (g axis) :
CONTROL IN ACCELERATION**



IN FLIGHT PRELIMINARY RESULTS

FLIGHT versus DROP TOWER TEST



$$\Phi \propto (Y2-Y1)$$

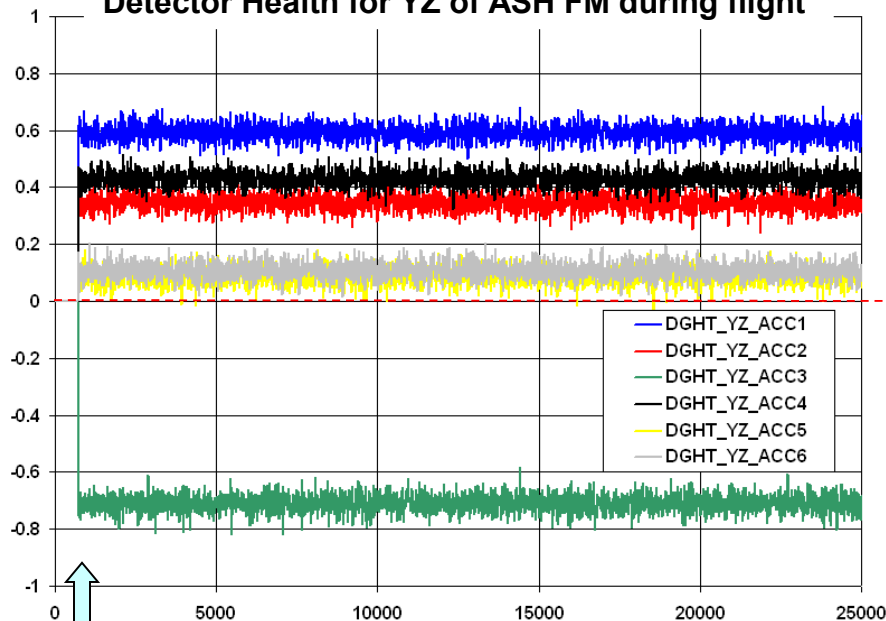
or

$$\Phi \propto (Z2-Z1)$$

Detector_Health Test:
 $|Y2-Y1 - Z1-Z2| < \text{Threshold}$

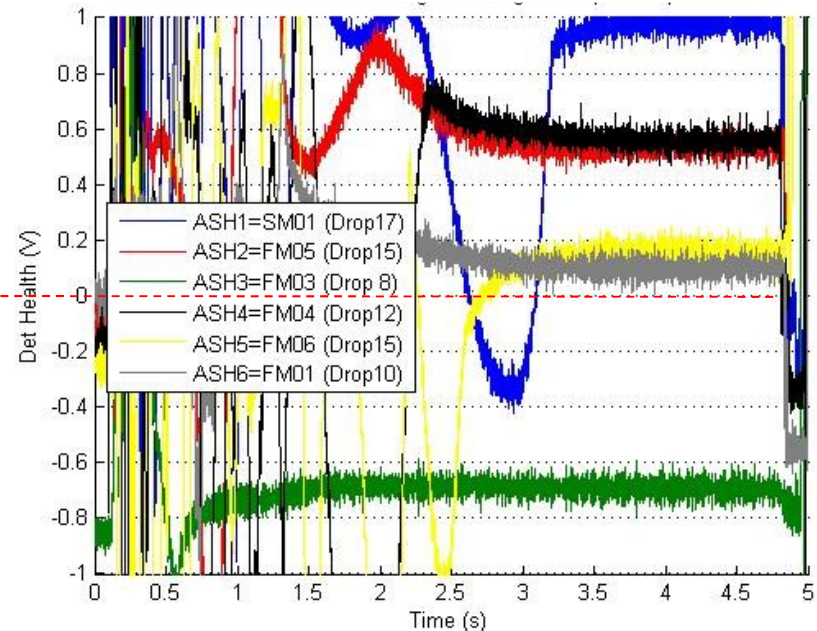
$$30\text{mV}_{\text{rms}} \Rightarrow 2\text{ nm}_{\text{rms}}$$

Detector Health for YZ of ASH FM during flight



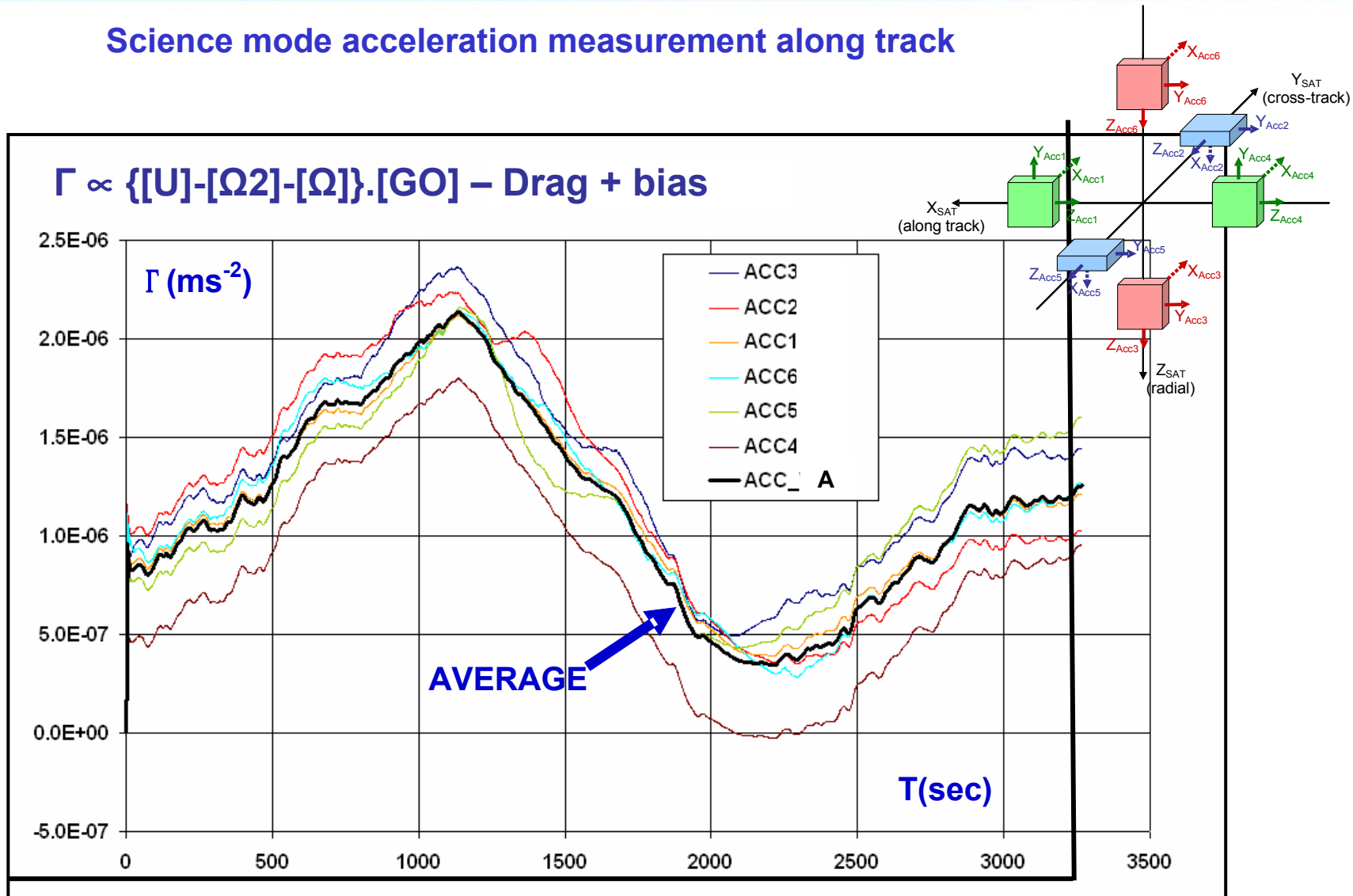
06/04/2009 9h26'40"

Detector Health for YZ of ASH FM during free fall



IN FLIGHT PRELIMINARY RESULTS

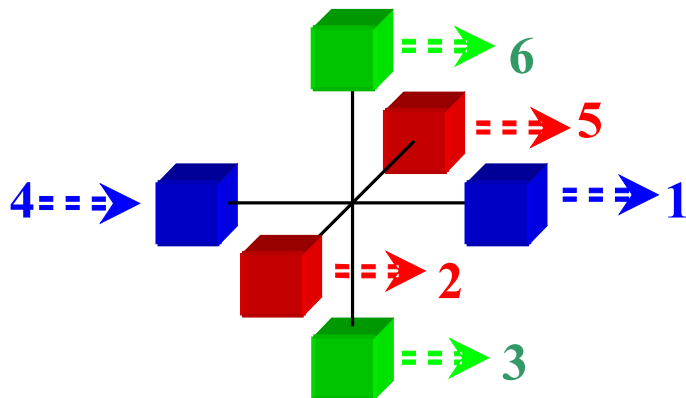
Science mode acceleration measurement along track



GOCE DRAG FREE PERFORMANCE

GOCE drag free performance, verified in commissioning phase (June 09) @270km

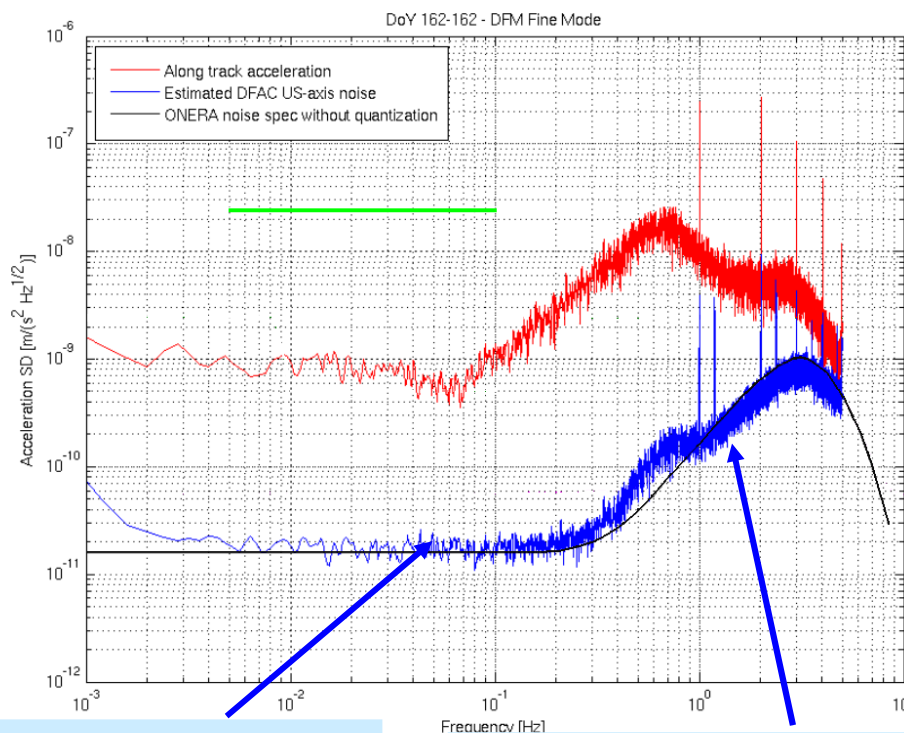
*GOCE: Space technology for the reference Earth gravity field determination,
A.Allasio, D.Muzi, B.Vinai, S.Cesare, G.Catastini, M.Bard, J.P. Marque, EUCASS 2009, Versailles France.*



The common mode of the two accelerometers voltages measures the non gravitational forces at the middle of their axis, located at the center of gravity of the satellite.

**DFAC noise estimated through
redundant acceleration measurements**

$$N = 2/\sqrt{3} [acc14x - (acc25x + acc36x)/2]$$



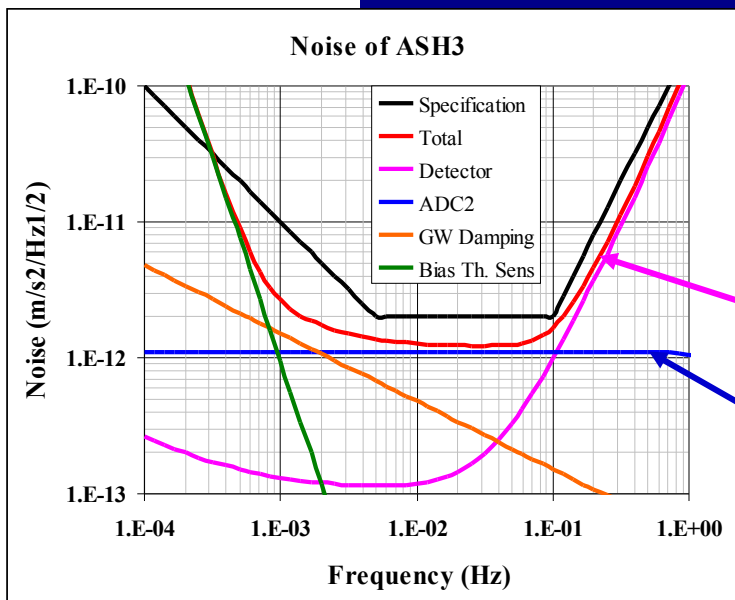
**Actuation noise (DAC + DVA)
main contributor
in [10 mHz – 100 mHz] MBW**

**Detection noise (detector + ADC1)
main contributor
in [0.5 Hz – 5 Hz] MBW**

PRESENT IN FLIGHT STATUS

Agreement between the in-orbit noise and the predicted error budget of the DFAC channel

- ❑ 6 accelerometers nominally operating
- ❑ Same behaviour of the 6 accelerometers
- ❑ Accelerometer control loops \Rightarrow OK
- ❑ W.r.t. foreseen Science channel Error Budget :
 - ❑ Detection noise \Rightarrow OK
 - ❑ Electrostatic actuation noise \Rightarrow OK



Validated by DFAC data

In flight test validation

GOCE ACCELEROMETER STATUS

- ❑ Mature Technology of the Mechanical sensor optimised in terms of design, manufacturing and integration processes to reach the necessary resistance and the stability (**< 8 μm over instrument life**) of the Accelerometer Reference Frame during the launch phase,
- ❑ Space quality electronics designed to reach an acceleration resolution better than **$2.0 \times 10^{-12} \text{ ms}^{-2} / \text{Hz}^{1/2}$** with a large measurement range of **$6.5 \times 10^{-6} \text{ ms}^{-2}$** ,
- ❑ Flexibility in the retrieval of the 6 Degrees of Freedom and redundancy with **8 electrode pairs** organised in **6 digital control loops per accelerometer**,
- ❑ **On Ground test ability** thanks to ASH geometrical configuration and dedicated EGSE with high voltage electronics and **Complete test plan to assess reliability**.
- ❑ **As inertial sensor of the DFAC system**, it provides highly accurate data to measure the 3 D acceleration of the gradiometer centre which is also the S/C CoG and it participates to fine attitude estimation.

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Return of Experience and towards next Gravity mission

1. PAST and PRESENT

STAR for CHAMP

- Γ_n : $3 \cdot 10^{-9} \text{ ms}^{-2} / \text{Hz}^{1/2}$
- Γ_{max} : 10^{-4} ms^{-2}

SuperSTAR for GRACE

- Γ_n : $1.0 \cdot 10^{-10} \text{ ms}^{-2} / \text{Hz}^{1/2}$
- Γ_{max} : $5 \cdot 10^{-5} \text{ ms}^{-2}$

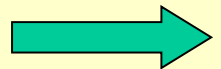
GRADIO for GOCE

- Γ_n : $2.0 \cdot 10^{-12} \text{ ms}^{-2} / \text{Hz}^{1/2}$
- Γ_{max} : $6 \cdot 10^{-6} \text{ ms}^{-2}$

Adjustability of the design w.r.t. performances:

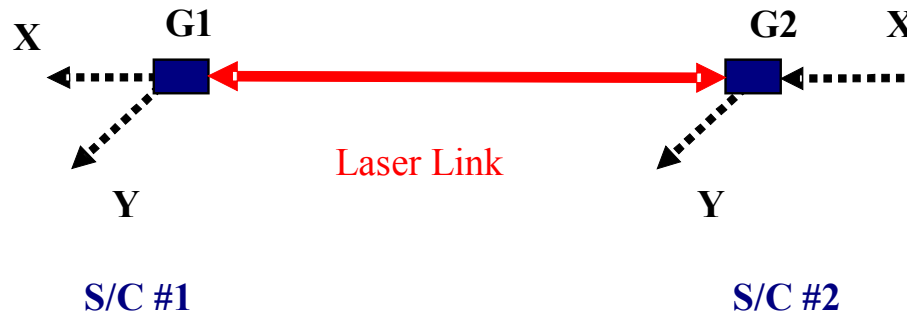
PM weight
gaps value
number of electrodes
loop design

Adaptability to the Satellite Configuration



Return of Experience and towards next Gravity mission

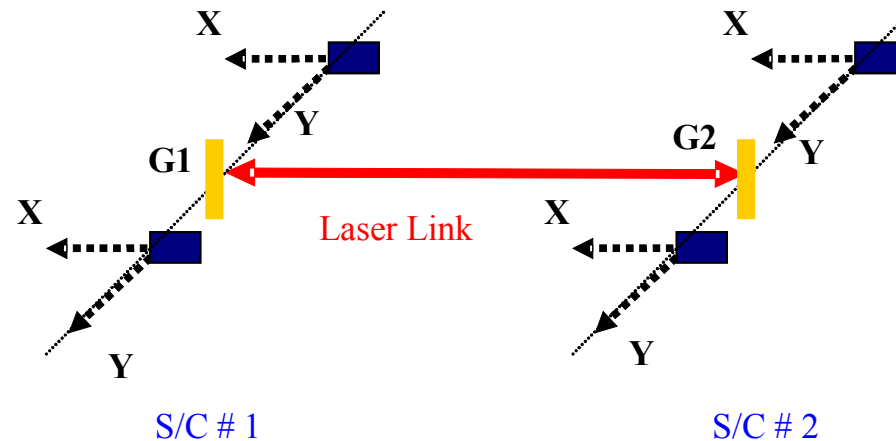
Case #1 Accelerometer at Center of Gravity as in CHAMP & GRACE



- Laser link better than microwave link
- To reflect link on proof-mass or not ?
- Drag free satellite without acceleration measurement :
 - optimized configuration of the inertial sensor:
 - larger gaps → reduced disturbances
 - but S/C motion w.r.t. mass : fluctuation of self gravity & CPD (GPB...)
 - other axes : coupling between axes ? linear and angular...

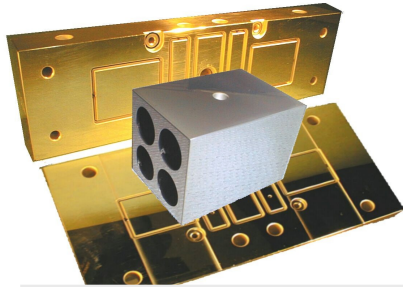
Return of Experience and towards next Gravity mission

Case #2 Accelerometer off-centered with Gradio pair at Center of Gravity (GOCE)



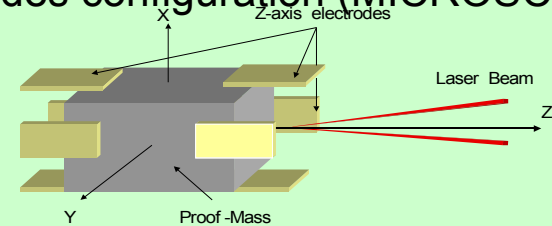
- S/C acceleration from common mode of each Gradio pair
- Accelerometers and no inertial sensors:
 - = drag free sensor + gradiometer sensor + accelerometer sensor for laser link
- S/C drag free satellite in all directions ?

Improvements for accurate future gravity mission

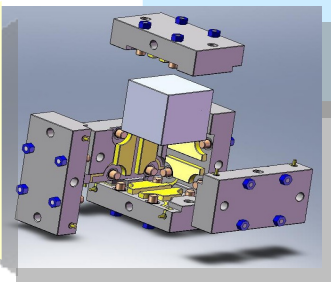
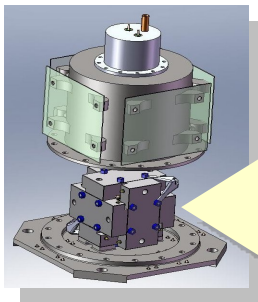


In the case of laser link between the SST satellites possibility to use the proof-mass itself as mirror for the interferometer.

Linear actuation (and capacitive sensing) through surface variation electrodes configuration (MICROSCOPE Conf.)



Optical position sensing using fiber laser interferometry techniques:
off-loop detection, very low back-action



Mass, volume and power budgets reduction:
MICROSTAR model
~1kg – 1 l – 1 W)

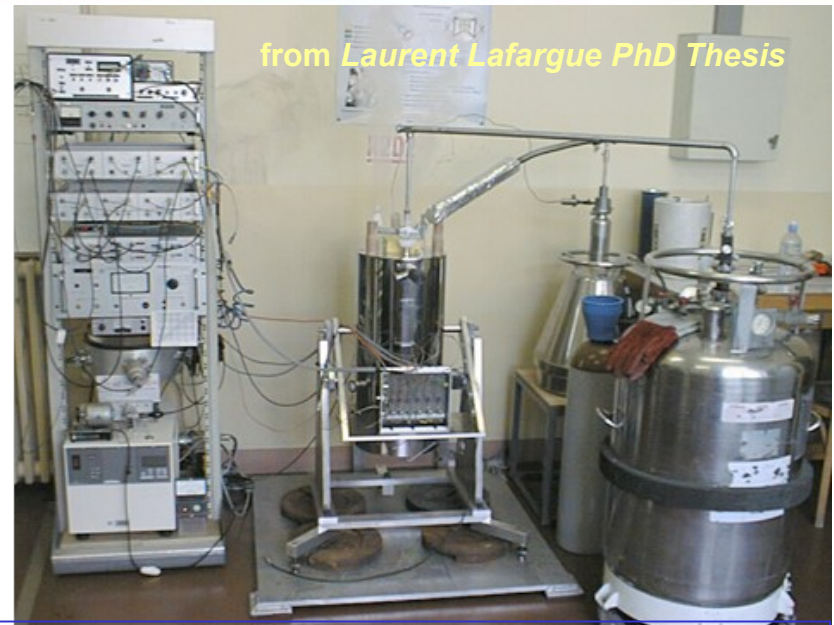
Atomic Interferometry based
accelerometer (ICE Project)

Cryogenic temperature instrument :
operation verified (*Laurent Laffargue PhD Thesis*), thermodynamic noise reduced by one order of magnitude, improvement of thermal and mechanical stabilities

Cryogenics accelerometers

Operation verified at Low Temperature:

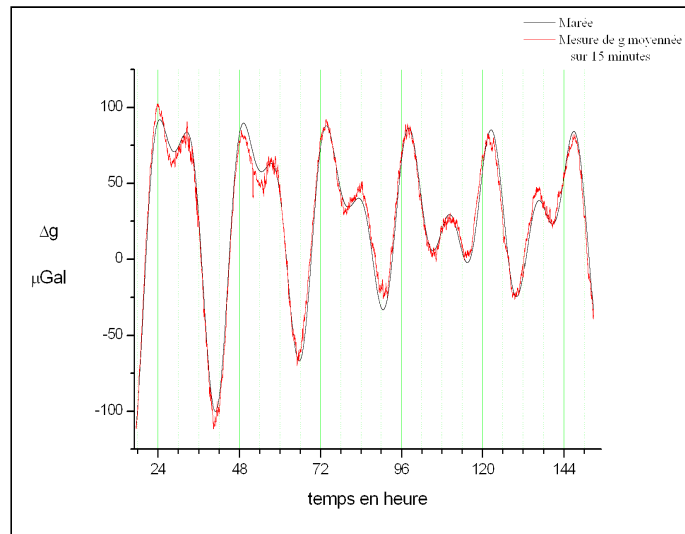
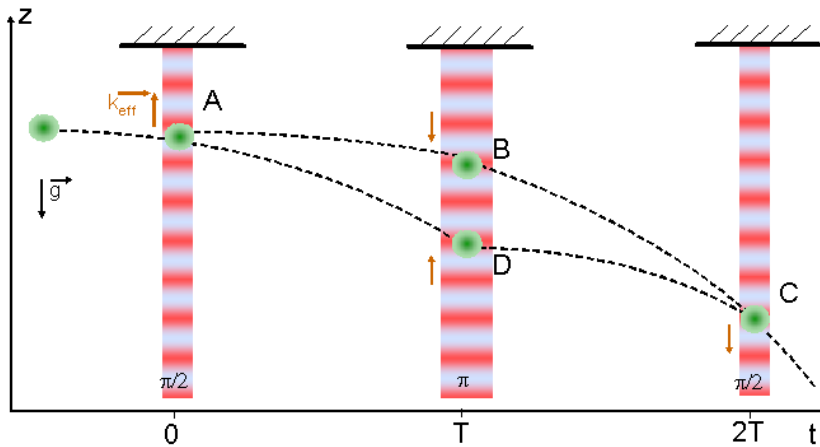
- Thermodynamic noise reduced by one order of magnitude,
- Improvement of thermal and mechanical stabilities



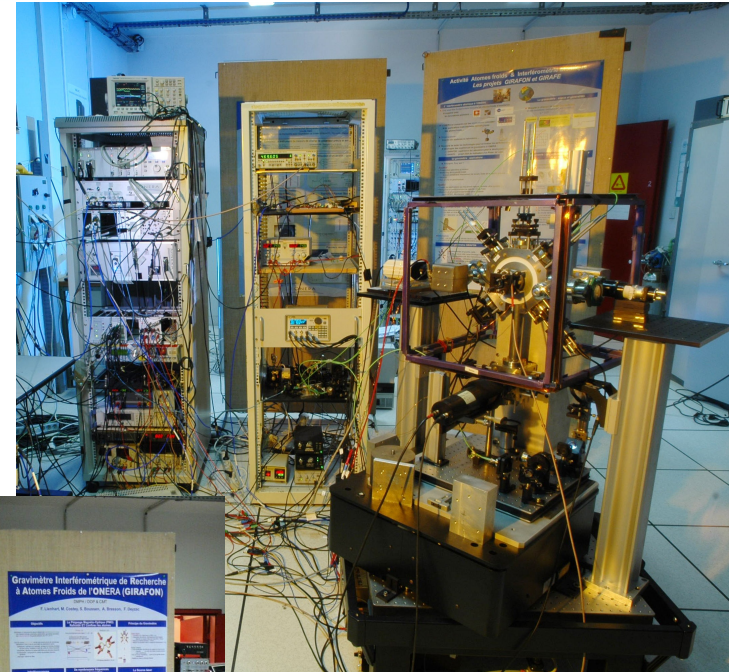
Superconductive :

- SQUID + Magnetic suspension :
 - no servo-control
 - necessity of electrostatic control and damping in addition
 - Electrostatic:
 - easy operation with room temperature electronics
 - reduced thermodynamic noise to be associated to active charge control
- expected gain in resolution up to 100 w.r.t. GOCE accelerometer range to be adapted

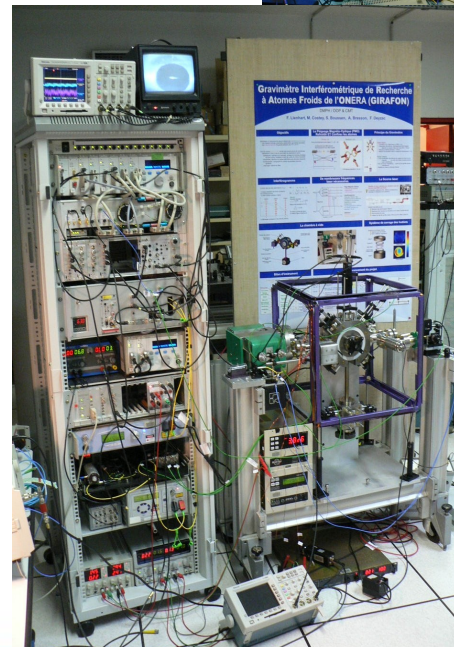
Atomic gravimeters also exist at Onera/DMPH !



Daily gravity field variation



Laboratory Setup



CONCLUSION

The 6 GOCE accelerometers are flight operational

**9 Electrostatic Accelerometers are presently working in orbit
(CHAMP,GRACE,GOCE)**

Robustness and Flexibility of the Design

Adjustability of the design w.r.t. performances

(PM weight, gaps value, number of electrodes,loop design....)

**Demonstration of drag compensation at low altitude
(ultra sensitive electrostatic accelerometer + ion propulsion)**

**Possibility to combine LL-SST with Gradiometry from
applications of accelerometers in GRACE and GOCE**



Thank you for your attention

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